

# Analysis of the asymmetry factor in photophoresis problems with uniform polychromatic fields and spherical absorbers

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## ABSTRACT

Photophoresis with plane waves and spherical absorbers is traditionally investigated by considering monochromatic radiation. The asymmetry factor, which is proportional to the photophoretic force, can be expressed in using the Mie theory in terms of an infinite sum of terms depending on the wavelength of the impinging light. Here, we analyze an extension of the photophoresis problem for a polychromatic uniform optical field. Such an analysis demands the introduction of a spectrum which describes the weight of each plane wave for the total, polychromatic field. We expect that this work can find applications or pave the way for more advanced investigations in diverse areas of physics and engineering such as optical levitation, optical trap displays and spectroscopy.

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## 1. Introduction

In all that follows, the problem consists of an opaque, non-radiative, non-volatile, non-magnetic spherical absorber of size parameter  $x = ka$  ( $a$  being its radius and  $k = 2\pi/\lambda$  being the wave number,  $\lambda$  is the wavelength) immersed in a gaseous fluid. Relatively to the host medium, it has a refractive index  $M = n_{sp} - im_{sp} = (\epsilon_{sp,r})^{1/2}$ . From an optical point of view, the particle is therefore completely described in terms of a relative electric permittivity  $\epsilon_{sp,r} = \epsilon'_r - i\epsilon''_r$ . Notice that a time-dependence factor of the form  $\exp(+i\omega t)$  is assumed, with  $\omega$  being the angular frequency, and that for polychromatic fields all these parameters might be frequency-dependent.

In photophoresis by monochromatic plane waves and low Reynolds number, the asymmetry factor, which first appeared in a work by Yalamov et al. (Yalamov et al., 1976), embeds in itself all the electromagnetic part of the problem. Here, following the conventions presented in Ref. (Ambrosio et al., 2022) and considering a  $+z$  propagating optical field, it shall be conveniently written in terms of the  $z$  component of the asymmetry vector  $\mathbf{r}_{as}$  (Ambrosio et al., 2022):

$$r_{as,z} = \hat{\mathbf{z}} \cdot \mathbf{r}_{as} = \frac{\epsilon_r''}{2} I_\lambda \mathcal{I}_z, \quad (1)$$

where  $I_\lambda = |E_0|^2/2\eta_m$  is the intensity of the wave,  $\eta_m$  is the intrinsic impedance of the medium and  $E_0$  is the electric field strength. In Eq. (1),  $\mathcal{I}_z$  is given in accordance with Eq. (17c) of Ref. (Ambrosio et al., 2022). In spherical coordinates  $(r, \theta, \phi)$ :

$$\mathcal{I}_z = s \frac{x}{a} \int_0^{2\pi} \int_0^\pi \int_0^a B_z(r, \theta, \phi) r^3 \cos \theta \sin \theta dr d\theta d\phi, \quad (2)$$

with  $B_z(r, \theta, \phi) = |\mathbf{E}_{int}(r, \theta, \phi)|^2/|E_0|^2$  being known as the dimensionless radiative intensity distribution function, source strength or normalized source function (Mackowski, 1989; Arnold et al., 1984; Greene et al., 1985). It must be emphasized that  $B_z$  is usually a fundamental parameter in photophoresis problems in the determination of the heat source function (the excitation term or inhomogeneous part of the heat transfer equation)  $Q(r, \theta, \phi) = -(1/2) \text{Re}[\nabla \cdot (\mathbf{E}_{int} \times \mathbf{H}_{int}^*)] = k\epsilon_r'' I_\lambda B_z(r, \theta, \phi)$  which is seen to be implicitly present in Eq. (1), since, from Eq. (2),  $x/a = k$ . Analytical solutions to Eq. (2) can be found, e.g., in Eq. (30b) of Ref. (Ambrosio et al., 2022) or, equivalently, in Eq. (8) of Ref. (Ambrosio, 2021), and read as:

$$\begin{aligned} \mathcal{I}_z \equiv \mathcal{I}_z(x, M, \omega) = & \frac{8\pi a^3}{|M|^2 x^3} \text{Im} \sum_{n=1}^{\infty} \left\{ \frac{n(n+2)}{M} (c_n^* c_{n+1} R_{n+1} + |\eta_r|^2 d_n^* d_{n+1} R_n) \right. \\ & \left. - \left[ \frac{n(n+2)}{n+1} (c_n c_{n+1}^* + |\eta_r|^2 d_n^* d_{n+1}) + \eta_r^* \frac{2n+1}{n(n+1)} c_n d_n^* \right] S_n \right\} \end{aligned} \quad (3)$$

In Eq. (3),  $c_n$  and  $d_n$  are the internal Mie coefficients and depend on the operating frequency, particle's geometry and electromagnetic properties (Gouesbet & Gréhan, 2017). Finally,  $\eta_r$  is the relative intrinsic impedance of the particle which, for a non-magnetic particle, can be written exclusively in terms of  $M$ , since  $\eta_r = 1/\sqrt{\epsilon_{sp,r}} = 1/M$ . Expressions for the quantities  $R_n$  and  $S_n$  can be found in Ref. (Mackowski, 1989), see Eqs. (59) and (61), with a correction for the sign in the latter, see for instance Eq. (14) and subsequent comments in Ref. (Ambrosio, 2020). For the convenience of the reader, they are here reproduced:

$$R_n = \frac{\text{Im} [M\psi_{n+1}(Mx) \psi_n^*(Mx)]}{\text{Im}[M^2]} \quad (4)$$

and

$$\begin{aligned} S_n = & -\frac{i}{2 \text{Im}[M^2]} \left\{ x (M|\psi_n(Mx)|^2 + M^*|\psi_{n+1}(Mx)|^2) \right. \\ & \left. - \left( M + 2(n+1) \frac{\text{Re}[M^2]}{M} \right) R_n + (2n+1) M^* R_{n+1} \right\} \end{aligned} \quad (5)$$

where  $\psi_n(z)$  is a Ricatti-Bessel function Gouesbet & Gréhan (2017).

The electric field  $\mathbf{E}_{int}$  inside the sphere is given in accordance with the Mie theory. Explicitly introducing the  $\omega$ -dependence, they can be written as:

$$E_{int,r} \equiv E_{int,r}(\mathbf{r}, \omega) = \sum_{n=1}^{\infty} E_0(\omega) (-i)^{n+1} (2n+1) c_n(\omega) \frac{\psi_n[k_{sp}(\omega)r]}{[k_{sp}(\omega)r]^2} P_n^1(\cos \theta) \cos \phi, \quad (6)$$

$$E_{int,\theta} \equiv E_{int,\theta}(\mathbf{r}, \omega) = \sum_{n=1}^{\infty} E_0(\omega) (-i)^{n+1} E_n \left[ c_n(\omega) \frac{\psi'_n[k_{sp}(\omega)r]}{k_{sp}(\omega)r} \tau_n^1(\cos \theta) - i \frac{k(\omega)}{k_{sp}(\omega)} d_n(\omega) \frac{\psi_n[k_{sp}(\omega)r]}{k_{sp}(\omega)r} \pi_n^1(\cos \theta) \right] \cos \phi, \quad (7)$$

$$E_{int,\phi} \equiv E_{int,\phi}(\mathbf{r}, \omega) = i \sum_{n=1}^{\infty} E_0(\omega) (-i)^{n+1} E_n \left[ i c_n(\omega) \frac{\psi'_n[k_{sp}(\omega)r]}{k_{sp}(\omega)r} \pi_n^1(\cos \theta) + \frac{k(\omega)}{k_{sp}(\omega)} d_n(\omega) \frac{\psi_n[k_{sp}(\omega)r]}{k_{sp}(\omega)r} \tau_n^1(\cos \theta) \right] \sin \phi. \quad (8)$$

In Eqs. (6)-(8),  $E_n = (2n+1)/[n(n+1)]$ ,  $k_{sp} = Mk$  is the wave number in the particle,  $P_n^1(\cos \theta)$  is an associated Legendre polynomial, and  $\pi_n^1(\cos \theta) = P_n^1(\cos \theta)/\sin \theta$  and  $\tau_n^1(\cos \theta) = dP_n^1(\cos \theta)/d\theta$  are generalized Legendre functions Gouesbet & Gréhan (2017). Once Eqs. (6)-(8) are replaced in Eq. (2), the asymmetry factor assumes the form provided by Mackowski in Ref. (Mackowski, 1989) under the time harmonic convention  $\exp(-i\omega t)$  or, alternatively, Eq. (8) of Ref. (Ambrosio, 2021) for  $\exp(+i\omega t)$ , with the coefficients  $g_n$  in this equation set equal to 1, in accordance with Eq. (3) above.

## 2. Extension to the polychromatic incidence

Let us now extend the previous analysis to polychromatic uniform fields constructed from superpositions of plane waves. To do so, remember that the instantaneous phasor form for the fields in Eqs. (6)-(8) is  $\mathbf{E}_{int} \exp(+i\omega t)$  and let us consider a polychromatic field of the form

$$\mathbf{X}_{int}(\mathbf{r}, t) = \int_{-\infty}^{\infty} \mathbf{X}_{int}(\mathbf{r}, \omega) e^{+i\omega t} d\omega, \quad (9)$$

with the inverse relation

$$\mathbf{X}_{int}(\mathbf{r}, \omega) = \frac{1}{2\pi} \int_{-\infty}^{\infty} \mathbf{X}_{int}(\mathbf{r}, t) e^{-i\omega t} dt, \quad (10)$$

In Eqs. (9) and (10),  $\mathbf{X} = \mathbf{E}$  or  $\mathbf{H}$ . As a particular case of Eqs. (9) and (10), for an optical field composed of red, green and blue plane waves (say, with  $\lambda = 632.8, 543.5$  and  $441.6$  nm, respectively):

$$\mathbf{E}_0(\omega) = \sum_{j=1}^3 \mathbf{E}_0(\omega_j) \delta(\omega - \omega_j), \quad (11)$$

with  $\mathbf{E}_0(\omega_j)$  corresponding to the electric field strength of the  $j$ -th plane wave. In Eq. (11),  $\delta(\cdot)$  is the Dirac delta function. Finally, Eq. (9) is further reduced for the instantaneous phasor form of a plane wave at a fixed angular frequency  $\omega_0$  by setting  $\mathbf{E}_0(\omega) = E_0 \delta(\omega - \omega_0)$ .

The effect of Eqs. (9) and (10) on the heat source function can be quantified as follows. First, we evaluate  $\mathbf{E}_{int} \times \mathbf{H}_{int}^*$ :

$$\begin{aligned} \mathbf{E}_{int} \times \mathbf{H}_{int}^* &= \int_{-\infty}^{\infty} \mathbf{E}_{int}(\mathbf{r}, \omega) e^{i\omega t} d\omega \times \int_{-\infty}^{\infty} \mathbf{H}_{int}^*(\mathbf{r}, \omega') e^{-i\omega' t} d\omega' \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \mathbf{E}_{int}(\mathbf{r}, \omega) \times \mathbf{H}_{int}^*(\mathbf{r}, \omega') e^{i(\omega - \omega')t} d\omega d\omega' \end{aligned} \quad (12)$$

Equation (12) is averaged over time using the fact that:

$$\lim_{T \rightarrow \infty} \frac{1}{T} \int_{-T/2}^{T/2} e^{i(\omega - \omega')t} dt = \begin{cases} 1, & \omega = \omega' \\ 0, & \text{otherwise} \end{cases} \quad (13)$$

thus leading to

$$\langle \mathbf{E}_{int}(\mathbf{r}, t) \times \mathbf{H}_{int}^*(\mathbf{r}, t) \rangle = \int_{-\infty}^{\infty} \mathbf{E}_{int}(\mathbf{r}, \omega) \times \mathbf{H}_{int}(\mathbf{r}, \omega) d\omega. \quad (14)$$

From Eqs. (12)-(14), the heat source function then becomes

$$\begin{aligned} Q &= -\frac{1}{2} \text{Re} [\nabla \cdot \langle \mathbf{E}_{int}(\mathbf{r}, t) \times \mathbf{H}_{int}^*(\mathbf{r}, t) \rangle] = -\frac{1}{2} \text{Re} \left[ \nabla \cdot \left( \int_{-\infty}^{\infty} \mathbf{E}_{int}(\mathbf{r}, \omega) \times \mathbf{H}_{int}^*(\mathbf{r}, \omega) d\omega \right) \right] \\ &= -\frac{1}{2} \text{Re} \left[ \int_{-\infty}^{\infty} \nabla \cdot (\mathbf{E}_{int}(\mathbf{r}, \omega) \times \mathbf{H}_{int}^*(\mathbf{r}, \omega)) d\omega \right] = \int_{-\infty}^{\infty} \frac{1}{2} \omega \epsilon_m \epsilon_r''(\omega) |\mathbf{E}_{int}(\mathbf{r}, \omega)|^2 d\omega \\ &= \int_{-\infty}^{\infty} Q(\mathbf{r}, \omega) d\omega \end{aligned} \quad (15)$$

For the multichromatic case given in Eq. (11), Eq. (12) becomes:

$$\begin{aligned}\mathbf{E}_{int} \times \mathbf{H}_{int}^* &= \sum_{j=1}^3 \mathbf{E}_{int}(\mathbf{r}, \omega_j) e^{i\omega_j t} \times \sum_{p=1}^3 \mathbf{H}_{int}(\mathbf{r}, \omega_p) e^{-i\omega_p t} \\ &= \sum_{j=1}^3 \sum_{p=1}^3 \mathbf{E}_{int}(\mathbf{r}, \omega_j) \times \mathbf{H}_{int}(\mathbf{r}, \omega_p) e^{i(\omega_j - \omega_p)t}.\end{aligned}\quad (16)$$

Equation (16) is again averaged over time. Instead of Eq. (13), one now uses the fact that:

$$\lim_{T \rightarrow \infty} \frac{1}{T} \int_{-T/2}^{T/2} e^{i(\omega_j - \omega_p)t} dt = \delta_{\omega_j, \omega_p}, \quad (17)$$

where  $\delta_{ij}$  is the Kronecker delta. This leads to

$$\langle \mathbf{E}_{int}(\mathbf{r}, t) \times \mathbf{H}_{int}^*(\mathbf{r}, t) \rangle = \sum_{j=1}^3 \mathbf{E}_{int}(\mathbf{r}, \omega_j) \times \mathbf{H}_{int}(\mathbf{r}, \omega_j) \quad (18)$$

Similarly, from Eqs. (16)-(18) the following expression is found for the heat source function:

$$\begin{aligned}Q &= -\frac{1}{2} \operatorname{Re} \left[ \nabla \cdot \left( \sum_{j=1}^3 \mathbf{E}_{int}(\mathbf{r}, \omega_j) \times \mathbf{H}_{int}^*(\mathbf{r}, \omega_j) \right) \right] = \sum_{j=1}^3 \frac{1}{2} \omega_j \epsilon_m \epsilon_r''(\omega_j) |\mathbf{E}_{int}(\mathbf{r}, \omega_j)|^2 \\ &= \sum_{j=1}^3 Q(\mathbf{r}, \omega_j)\end{aligned}\quad (19)$$

It is then seen that Eq. (15) implies on  $\mathcal{I}_z$  assuming the form

$$\mathcal{I}_z = \int_{-\infty}^{\infty} \mathcal{I}_z(x, M, \omega) d\omega, \quad (20)$$

while Eq. (19) leads to

$$\mathcal{I}_z = \sum_{j=1}^3 \mathcal{I}_z(x, M, \omega_j). \quad (21)$$

In Eqs. (20) and (21), the integrand and the terms in the summation are calculated through Eq. (3) for an arbitrary frequency  $\omega$  or for a discrete set of frequencies  $\omega_j$ . Notice further that it is only possible to derive Eqs. (20) and (21) through a time-averaging process of the optical fields, since such a process allows one to rule out the interference terms observed in Eqs. (12) and (16). The same process reveals that  $r_{as,z}$  in Eq. (1) assumes the same form of Eqs. (20) and (21).

### 3. Examples

As a first example of calculation of  $r_{as,z}$ , let us consider Eq. (21) and assume a polystyrene sphere whose real and imaginary parts of its refractive index with respect to air will be assumed to take values retrieved from Fig. 2 of Ref. (Velazco-Roa & Thennadil, 2007). The host medium is air ( $\eta_m \approx 377 \Omega$ ), and three wavelengths of  $\lambda_1 = 450$  nm,  $\lambda_2 = 700$  nm and  $\lambda_3 = 950$  nm are considered for the calculations. The field strengths are normalized, such that  $E_0(\omega_j) = 1, \forall j$ .

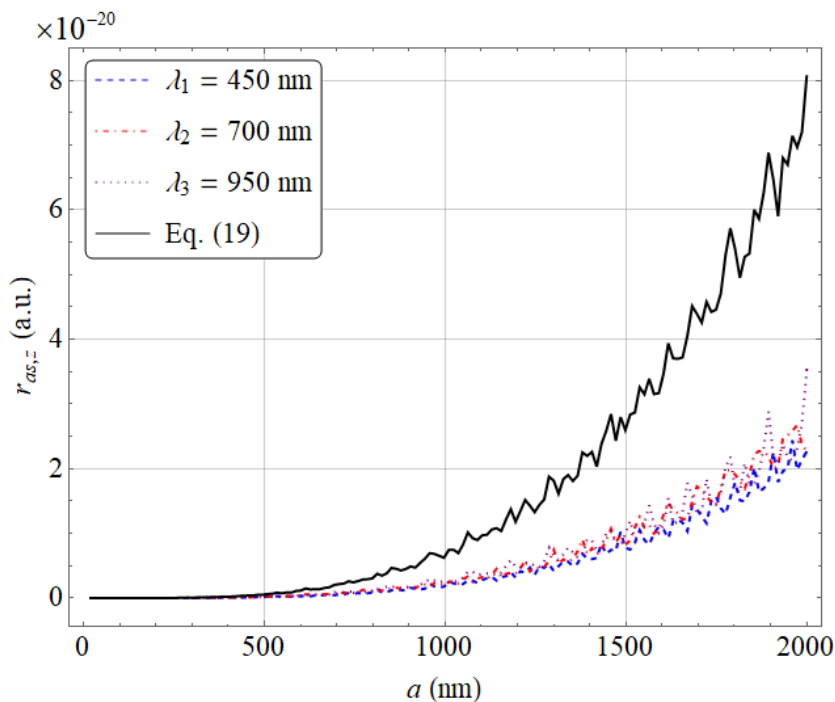
From the data presented for the Mie curves in Fig. 2 of Ref. (Velazco-Roa & Thennadil, 2007), one finds approximate values of  $M_1 = 1.640 - i2.205 \times 10^{-3}$ ,  $M_2 = 1.588 - i3.453 \times 10^{-3}$  and  $M_3 = 1.564 - i4.224 \times 10^{-3}$  for the wavelengths  $\lambda_1$ ,  $\lambda_2$  and  $\lambda_3$ , respectively. This means that such particles are weak absorbers. Figure 1 shows  $r_{as,z}$  as a function of the radius of the particle ( $20 \text{ nm} \leq a \leq 2 \mu\text{m}$ ) for each wavelength, together with the total  $r_{as,z}$ . For the entire range of  $a$ , it is seen that  $r_{as,z} > 0$ , which corresponds to a negative photophoretic force. Notice that the total  $r_{as,z}$  mimics the behavior of the individual curves, whose overall behavior are quite similar.

As a second example, we now assume a gold plasmonic particle. Such a spherical scatterer is highly absorbent, and its refractive index changes significantly over the range of optical frequencies [10]. For instance, according to data retrieved from Ref. (Magnozzi et al., 2018) (see also data available at <https://refractiveindex.info/?shelf=main&book=Au&page=Magnozzi-25C>) and for the same wavelengths as before,  $M_1 = 1.4783 - i1.86$ ,  $M_2 = 0.12309 - i4.1567$  and  $M_3 = 0.16524 - i6.2423$ . Figure 2 reveals the behavior of  $r_{as,z}$  for the same radii as before. In contrast with Fig. 1,  $r_{as,z} < 0$ , which corresponds to a positive photophoretic force.

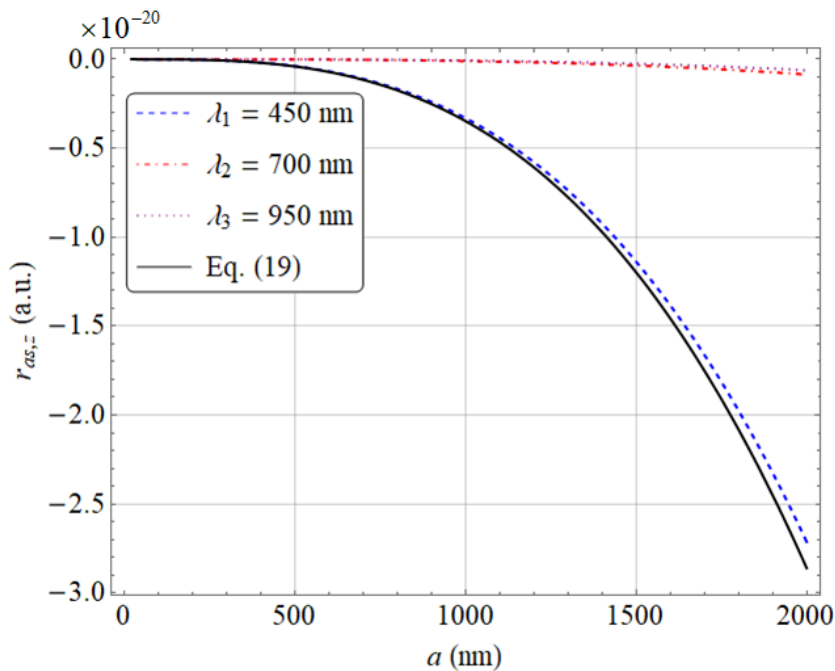
As a final example, in Fig. 3 we see the expected  $r_{as,z}$  as a function of  $a$  for Gallium arsenide (Pap-tryfonos et al., 2021), which is a fundamental material in optoelectronics (see also data available at <https://refractiveindex.info/?shelf=main&book=GaAs&page=Paptryfonos>). Here,  $\lambda_1 = 500$  nm,  $\lambda_2 = 700$  nm and  $\lambda_3 = 900$  nm, for which  $M_1 = 4.2230 - i0.45325$ ,  $M_2 = 3.7223 - i0.12964$  and  $M_3 = 3.5325 - i1.9044 \times 10^{-4}$ . In Fig. 3(a) we have weighted the field strengths such that  $E_0(\omega_1) = E_0(\omega_2) = E_0(\omega_3) = 1$ , while in Fig. 3(b) we have set  $E_0(\omega_1) = E_0(\omega_2) = E_0(\omega_3)/1000 = 1$ . It becomes clear that the total value of  $r_{as,z}$  shifts from positive to negative by simply considering different field strengths. Therefore, in the more general case where  $E_0(\omega)$  represents a spectrum over, say, the visible or any other range, the overall behavior for the asymmetry factor can be quite distinguished from that arising from a monochromatic incidence.

### 4. conclusion

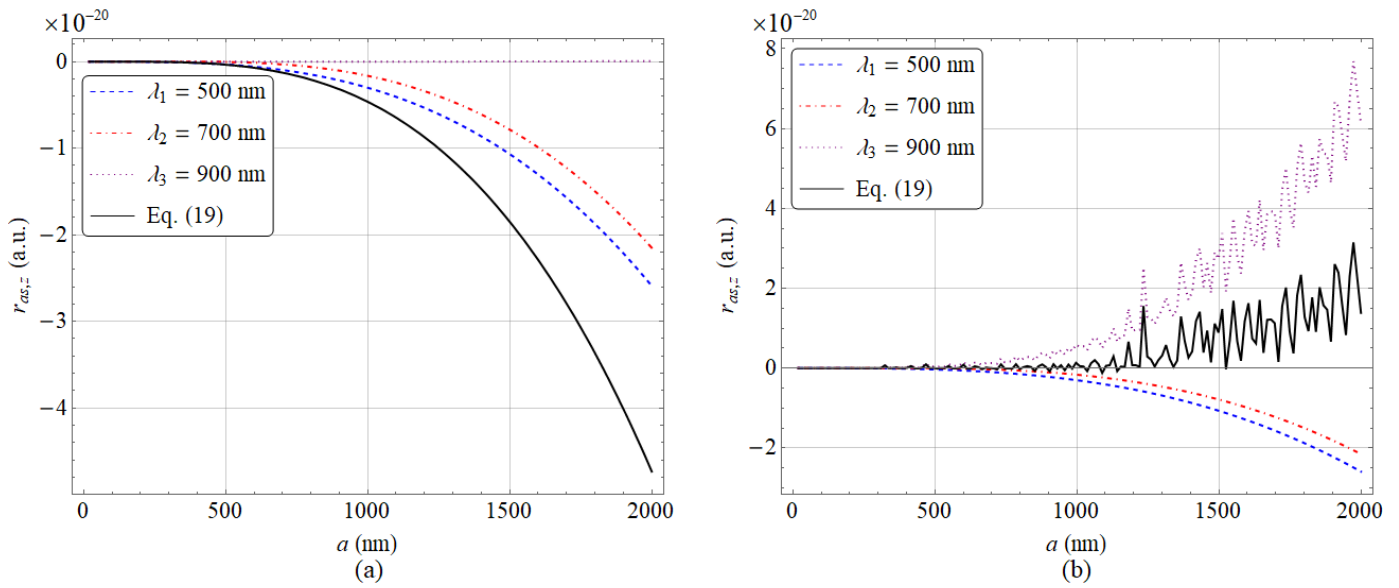
In this work we have presented, for the first time to the best of the authors' knowledge, an approach to investigate photophoresis with polychromatic uniform light. The theoretical analysis indicates that the asymmetry factor, after a time averaging process, can be simply written as



**Figure 1.**  $r_{as,z}$  as a function of  $a$  for a multichromatic light field of wavelengths  $\lambda_1 = 450$  nm,  $\lambda_2 = 700$  nm and  $\lambda_3 = 950$  nm. The particle is made of polystyrene, with  $M_1 = 1.640 - i2.205 \times 10^{-3}$ ,  $M_2 = 1.588 - i3.453 \times 10^{-3}$  and  $M_3 = 1.564 - i4.224 \times 10^{-3}$  and  $E_0(\omega_j) = 1, \forall j$ .



**Figure 2.** Same as Fig. 1, now for a gold particle with  $M_1 = 1.4783 - i1.86$ ,  $M_2 = 0.12309 - i4.1567$  and  $M_3 = 0.16524 - i6.2423$ .



**Figure 3.**  $r_{as,z}$  as a function of  $a$  for a multichromatic light field of wavelengths  $\lambda_1 = 500$  nm,  $\lambda_2 = 700$  nm and  $\lambda_3 = 900$  nm. The particle is made of gallium arsenide (GaAs), with  $M_1 = 4.2230 - i0.45325$ ,  $M_2 = 3.7223 - i0.12964$  and  $M_3 = 3.5325 - i1.9044 \times 10^{-4}$ . (a)  $E_0(\omega_1) = E_0(\omega_2) = E_0(\omega_3) = 1$ . (b)  $E_0(\omega_1) = E_0(\omega_2) = E_0(\omega_3)/1000 = 1$ .

weighted (continuous or discrete) sum of the asymmetry factor for monochromatic plane waves, the field strengths now playing the role, in the case of a continuous dependence on angular frequency, of a spectrum. The study here proposed is still limited to uniform light and non-magnetic materials, and it would certainly be of interest to extend it in order to incorporate both arbitrary-shaped waves [by having resource, for instance, to the polychromatic generalized Lorenz-Mie theory proposed in 2022 by Ambrosio et al. (2024)] and arbitrary-index particles like negative index, magnetodielectric and other conventional or exotic materials carrying also a magnetic response. This is currently in progress.

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